Comparing Solar-Flare Acceleration of >~20 MeV Protons and Electrons Above Various Energies

Albert Y. Shih¹, R. P. Lin², B. R. Dennis¹

¹NASA Goddard Space Flight Center ²SSL, UC Berkeley



Composite photon spectrum



Comparing ions and electrons

- Do all flares accelerate ions and electrons?
- Flare-acceleration models do not typically predict a constrained ratio of ion and electron acceleration
- Use ion-associated and electron-associated emissions as measures of particle acceleration
- Observations from *RHESSI* and *SMM*/GRS

Versus >300 keV electrons



- Direct proportionality
- Dotted lines are factors of 2 from the best-fit line
- Some spread is due to incomplete coverage (triangles)
- Almost all flares fall within 1 σ of spread

Versus thermal emission



- GOES class (emission from hot plasma) as a measure of flare size
- Gray region has not been systematically searched
- Direct proportionality above a threshold?
- Below threshold: excess heating

Versus >50 keV electrons



- Subset of the *RHESSI* flares that are easier to analyze
- Many flares show comparable correlation as with >300 keV fluence
- Five flares appear to deviate significantly: excess >50 keV?

Flare count spectra: 50–600 keV



Modified >50 keV correlation



- Now excluding the contribution of soft, low-energy bremsstrahlung
- Extrapolated the high-energy power law down to 50 keV
- The flares with good statistics correlate much better

Conclusions

- >~20 MeV ions and >300 keV electrons are proportionally accelerated over >3 orders of magnitude
- >~20 MeV ions and >50 keV electrons are not necessarily proportionally accelerated because of soft, low-energy components (<~ 150–300 keV)
- Excess thermal emission is possibly associated with this low-energy component